



Interannual variability
of UTLS over the
GBM river basin

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Interannual variability of upper tropospheric and lower stratospheric (UTLS) region over Ganges–Brahmaputra–Meghna basin based on COSMIC GNSS RO data

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Abstract

Poor reliability of radiosonde observational networks across South Asia imposes serious challenges in understanding climate variability and thermodynamic structure of the upper-tropospheric and lower-stratospheric (UTLS) region. The Constellation Observing System for Meteorology, Ionosphere, and Climate (COSMIC) mission launched in April 2006 have overcome many observational limitations that are inherent in conventional atmospheric sounding instruments. This study investigated the interannual variability of the UTLS region over the Ganges–Brahmaputra–Meghna (GBM) basin based on COSMIC radio occultation (RO) data from August 2006 to December 2013. Detailed comparisons were also made with various different radiosonde types and Numerical Weather Prediction (NWP) products. The results indicated that Indian Meteorological Department (IMD) radiosondes performed poorly despite upgrading to newer techniques. ShangE (of China) sonde showed the best agreement with COSMIC RO data with a mean temperature difference of -0.06°C and a standard deviation of 1.44°C while the older version (ShangM) indicated a cold bias of 0.61°C in the UTLS region. The inter-annual variability of temperature in the UTLS region based on COSMIC RO data indicated a clear pattern of El Niño Southern Oscillation (ENSO) and Indian Ocean Dipole (IOD) while the stratospheric temperature anomalies reflected all three major Sudden Stratospheric Warming (SSW) events since 2007. The mean tropopause temperature varied from -70 to -80°C , with an average height of about 15.5 to 16.3 km from winter to summer, indicating a pronounced annual cycle. The annual amplitudes of tropopause were found to be in the order of $0-6^{\circ}\text{C}$ and $0-1.5$ km from tropical south to subtropical north. The anomalies of tropopause temperature and height exhibited the patterns of ENSO and IOD exceptionally well with a correlation of 0.65 and -0.52 , respectively. The temperature data from Modern-Era Retrospective Analysis for Research Application (MERRA) agreed very well with COSMIC RO data.

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1 Introduction

The interannual variability and change in the vertical structure of the upper-tropospheric/lower-stratospheric (UTLS) region (400–30 hPa) provides key insights into the underlying causes of global and regional aspects of climate change and is closely associated with the surface temperature fluctuations (Karl et al., 2006; Bindoff et al., 2013). The tropopause marks the separation between the two boundary layers in the UTLS region and is a key indicator of global climate change (Santer et al., 2003; Sausen and Santer, 2003) due to its role in tropospheric–stratospheric exchange processes. Observational evidences from balloon-borne radiosondes since the late 1950s and satellite-based measurements since 1979 suggest that the troposphere has warmed considerably over the past decades with substantial cooling in the lower stratosphere (Karl et al., 2006; Bindoff et al., 2013; Lott et al., 2013; Thorne et al., 2013). Much of this temperature changes has been attributed to the anthropogenic emissions of the well-mixed greenhouse gases.

While there is increasing evidence in the recent amplification of global and regional vertical structure of the troposphere and the lower stratosphere (see, e.g., Bindoff et al., 2013), large observational uncertainties exist especially in the UTLS region due to limited observations. As a result, trends in upper air temperatures might be of limited quality, thus, affecting the climate change attribution studies (Seidel et al., 2011; Steiner et al., 2011). Radiosondes can hardly reach up to 50 hPa (~ 20.5 km) as most of them tend to burst out in the upper tropospheric region. Other issues include instrumental changes over time and poor spatial coverage. Satellite-based Microwave Sounding Units (MSUs) used to retrieve atmospheric information in the UTLS region do not provide adequate vertical resolution and therefore, might indicate systematic biases (Seidel et al., 2011; Bindoff et al., 2013). For example, previous studies reported anomalously large warm biases in radiosonde observations in the South Asian region, particularly over India (see, e.g., Kumar et al., 2010; Sun et al., 2010). In order to improve on these limitations, a more robust space-based technique known as

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the Global Navigation Satellite System (GNSS) radio occultation (RO) (e.g., Melbourne et al., 1994; Ware et al., 1996; Kursinski et al., 1997; Awange, 2012) has emerged as an important climate monitoring system over the past decade (see, e.g., Khandu et al., 2011; Steiner et al., 2013).

GNSS RO technique utilises the time delay information of the occulted GNSS signals, which passes through the atmosphere and is received by Low Earth Orbiting (LEO) satellites. The primary observable are GNSS phase path and signal amplitude, which can be subsequently converted into atmospheric profiles (of e.g., refractivity, temperature) using the assumption of spherical geometry of refractivity (e.g., Melbourne et al., 1994; Rocken et al., 1997). Nowadays, the retrieved bending angles can be directly applied in climate- and weather-related studies, thereby introducing less uncertainty in the observation system (see, e.g., Lewis, 2009). The use of GNSS RO data is advantageous since it provides 24 h global coverage, high vertical resolution, and highly accurate profiles of the upper-tropospheric/lower-stratospheric (UTLS) region (e.g., Sun et al., 2010; Khandu et al., 2011). Therefore, it has continuously been used to analyse the recent variability in the vertical structure in the UTLS region both globally (e.g., Schmidt et al., 2008, 2010) and regionally (e.g., Rao et al., 2009; Fu, 2011; Khandu et al., 2011). The number of RO profiles have increased substantially over the past years with the launch of several GNSS RO missions enabling wider applications in regional studies (see, e.g., Anthes, 2011). For instance, the joint Taiwan-US mission Constellation Observing System for Meteorology, Ionosphere, and Climate (COSMIC)/FORMOSA Satellite Mission 3 (COSMIC/FORMOSAT-3, hereafter, COSMIC) (Anthes et al., 2008) has recorded about 1500 RO soundings per day globally with 70–90 % of the soundings reaching within one km of the Earth's surface.

While many studies have been conducted to re-assess the global as well as regional lower atmospheric structure using the RO profiles, including their accuracies (Rocken et al., 1997; Kuo et al., 2005; Schmidt et al., 2010; Sun et al., 2010; Khandu et al., 2011; Kishore et al., 2011; Danzer et al., 2014), only few such assessments have been carried out over the Indian sub-continent (e.g., Rao et al., 2009). The Ganges–

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Brahmaputra–Meghna (GBM) basin in South Asia offers an interesting view on the regional perspective of the variability of UTLS because of its hydrological importance and intense activity of Indian monsoon from June to September. Previous studies have suggested that tropopause heights could be used as an additional indicator of Indian monsoon due to its high correlation with rainfall in May (Kulkarni and Verma, 1993). The rising mountains along the southern foothills of the Himalayas act as barriers to the Indian monsoon generating deep convection along the Himalayan fronts, which may affect the spatial variability of tropopause parameters such as temperature and height. The growing population and expanding industrialisation in and around the GBM river basin are the major source of atmospheric pollution and greenhouse gases (GHGs), and a key modulator of the UTLS region including the tropopause (Gautam et al., 2009; Lau et al., 2009). Thus, long-term self-calibrated GNSS RO data from various LEO missions will help to provide valuable information on the regional climate change in future as indicated by several studies (see, e.g., Schmidt et al., 2008; Steiner et al., 2011, 2013).

This study investigates the interannual variability of temperature in the UTLS region and tropopause parameters (temperatures and heights) over the GBM river basin for the period August 2006 to December 2013 (89 months) using the GNSS RO profiles from the COSMIC RO mission. Further, COSMIC RO data are compared with Modern-Era Retrospective Analysis for Research Application (MERRA, Rienecker et al., 2008) to assess the interannual variability of temperature of the UTLS region. It should be pointed out that no specific study involving MERRA products have been carried over the region. Three specific objectives are outlined as follows; (i) intercompare COSMIC RO retrieved profiles with radiosonde observations and reanalyses products, with specific focus on the upgraded radiosondes over India, (ii) analyse the interannual variability of UTLS temperature (400–30 hPa or 7.5–24.0 km), and (iii) assess the interannual variability of tropopause over the GBM river basin. The use of retrieved parameters such as refractivity, temperature, and water vapour from COSMIC RO data accounts for the overall error budget of the GNSS RO technique.

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The remainder of the study is organised as follows. In Sect. 2, the study region is presented. This is followed in Sect. 3 by the description of datasets and methods used, with the interpolation algorithm used to interpolate the point-based COSMIC RO datasets presented in the Appendix. Section 4 discussed the results, and the study concluded in Sect. 5.

2 Ganges–Brahmaputra–Megha (GBM) basin

The GBM river basin in South Asia is a combination of three medium to large river basins, namely, Ganges basin (907 000 km²), Brahmaputra basin (583 000 km²) and the Meghna basin (65 000 km²) (Chowdhury and Ward, 2004). This transboundary river basin with an elevation range from the sea level to more than 8000 m is shared by 5 countries (India (64 %), China (18 %), Nepal (9 %), Bangladesh (7 %) and Bhutan (3 %), see, Fig. 1). The GBM river basin with a total surface area of approximately 1.75 million km² features distinct climatic characteristics owing to its diverse climate and other factors such as high topographic variations, the Indian Monsoon, and its interaction with large scale circulations (e.g., Chowdhury, 2003). For instance, Ganges basin is generally characterised by low precipitation while Brahmaputra and Meghna basins are characterised by high rainfall amount (Mirza et al., 1998). The Himalayan fronts (e.g., Meghalayan Plateau) act as an immediate barrier to the summer monsoonal flow and are usually characterized by pronounced rainfall along the Himalayan fronts and across the southern foothills (Barros et al., 2004).

The atmospheric conditions over the basin are largely controlled by the monsoonal circulation during summer (e.g., Kripalani et al., 2007), which is often modulated by global and regional large-scale climate variabilities such as El Niño Southern Oscillation (ENSO) and Indian Ocean Dipole (IOD) (e.g., Chowdhury, 2003; Ashok and Saji, 2007). The impact of global warming and regional climate change arising from increasing population and rapid industrial and agricultural activities, and landuse changes

across the basin may have resulted in a warmer troposphere over the years (Gautam et al., 2009; Lau et al., 2009).

3 Data and methods

3.1 FORMOSAT/COSMIC RO data

5 COSMIC Level 2 RO data analysed at the COSMIC Data Analysis and Archive Center (CDAAC) at the University Corporation for Atmospheric Research (UCAR) for the period August 2006 to December 2013 was used in this study. CDAAC maintains an archive of RO profiles from all previous and existing RO missions such as COSMIC, CHALLENGING Minisatellite Payload (CHAMP), Gravity Recovery and Climate Experiment (GRACE), etc. COSMIC is a six-satellite mission launched on the 14 April 2006, with the goal of using GNSS RO data for various atmospheric and weather-related applications (Anthes et al., 2008). With an average sounding of around 1500–2000 profiles per day (at ~ 100 m to 1 km vertical resolution and ~ 300 km horizontal resolution), COSMIC has become a highly successful RO mission with high positive impacts especially on the operational weather forecasts (see e.g., Anthes et al., 2008; Anthes, 15 2011) and global atmospheric studies (see e.g., Foelsche et al., 2008; Schmidt et al., 2010).

The refractivity (N) and temperature (T) profiles retrieved from raw GNSS signals are referred to as the dry profiles. The raw refractivity profiles are separately processed using a 1-D-Var (one-dimensional variational) analysis with background information from numerical weather forecasts to provide accurate estimates of temperature and humidity in the troposphere. These profiles are referred as the wet profiles and contains information on refractivity, temperature, and water vapour pressure (or relative humidity). CDAAC uses a coarse grid background data from the European Centre for Medium-Range Weather Forecasts (ECMWF) re-analysis (see, http://cdaac-www.cosmic.ucar.edu/cdaac/products.html). The wet and dry profiles mainly

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differ in the lower troposphere due to presence of water vapour but are highly accurate between 8 and 20 km (see, Anthes et al., 2008). The a priori information (i.e., ECMWF) is minimally influenced by radiosonde observations in the GBM basin as they were not included in ECMWF (Kumar et al., 2010).

The GBM basin received a total of 35 776 profiles from April 2006 to December 2013, at an average of ~ 388 profiles per month (see, Fig. 2a). Figure 2b shows the distribution of COSMIC RO data points at various altitude levels indicating that COSMIC data is able to penetrate deep into the lower troposphere with more than 56 % of the profiles reaching at least 850 hPa (~ 1.5 km above MSL). Figure 3a–d shows the spatial distribution of COSMIC RO data at 850 hPa (~ 1.5 km), 700 hPa (~ 3.1 km), 500 hPa (~ 5.8 km), and 400 hPa (~ 7.5 km), which indicated generally a good correlation with the surface topography of the basin (Fig. 1). The near-complete coverage of the profile data can be seen at 400 hPa (~ 7.5 km) corresponding to the elevation of Himalayas, with Mt. Everest (8848 m) being the highest.

3.2 Radiosonde data

CDAAC also maintains radiosonde records from around the globe, which are collocated with all the RO missions (e.g., COSMIC, CHAMP, GRACE). The radiosonde profiles are extracted from National Center for Atmospheric Research (NCAR) mass store, which is comprehensively described in Sun et al. (2010) including their sources and quality control procedures. There are 24 operational or synoptic radiosonde stations within or around the GBM basin, with most of them operated by the Indian Meteorological Department (IMD) (see, Fig. 1). Three radiosonde stations are located in southern China (or the upper Brahmaputra basin), three are located in southern Bangladesh and the rest (18) of them are located inside the Indian territory. Based on the country of location, these radiosondes differ in their sensor types. The details of the radiosondes are provided in Table 1. The accuracy of IMD radiosondes have been a concern for many years due to their poor performance (see, e.g., Das Gupta et al., 2005; Kumar et al., 2010; Sun et al., 2010) and has undergone major upgrades in the past 5–6 years.

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which are integrated forward in time using a finite-volume dynamics to produce a consistent set of spatio-temporal meteorological and climatic variables since the beginning of the satellite-era (i.e., 1979). An improved land-process model known as the Catchment Land Model has been integrated into the GCM to improve the global hydrological cycle (Rienecker et al., 2008). GEOS-5 is run at a horizontal resolution of $1/2^\circ \times 2/3^\circ$ (or $\sim 50\text{ km} \times 70\text{ km}$) and has 72 vertical layers extending from the surface through to the stratosphere. Atmospheric variables (e.g., temperature, humidity) are produced at various temporal scales ranging from 3 hourly at $1.5^\circ \times 1.5^\circ$ (or $\sim 150\text{ km} \times 150\text{ km}$ spatial resolution), to monthly scales at the nominal horizontal resolution.

GEOS-5 uses radiosondes as the predominant source of information in the model atmosphere, which are further augmented by other atmospheric sensors such as dropsondes, pilot balloons, and raw radiance measurements from satellites (e.g., Advanced MSU) that are weighted according to their observational error variances (see, Rienecker et al., 2008, for details). The variables (e.g., temperature, humidity) are reported at 11 pressure levels for the UTLS region between 400 hPa ($\sim 7.5\text{ km}$) to 30 hPa ($\sim 24.0\text{ km}$). While radiosondes are sparse in the GBM basin (see, Fig. 1), the MERRA reanalyses products serve as complementary data source in the lower atmosphere in addition to the existing reanalyses products (e.g., ERA-Interim). Here, MERRA temperature in the UTLS region and tropopause parameters (temperatures and heights) were compared with those of COSMIC RO data over the GBM basin for the period August 2006 to December 2013.

3.4 Other datasets

240 collocated profiles, each from two operational forecast products maintained by CDAAC were used as additional data to ascertain the reliability of radiosonde observations over the GBM river basin. The two operational products used are (a) Global Forecast Model (GFS) (Kalnay et al., 1996) from National Centers for Environmental Prediction (NCEP)/NOAA, and (b) ERA-Interim (Dee et al., 2011) from European Centre for Medium-Range Weather Forecasts (ECMWF). Both forecast products have

been widely used globally to evaluate global as well as regional observations and climate models (e.g., Lott et al., 2013; Thorne et al., 2013; Vergados et al., 2015) and are largely independent of the radiosonde datasets over the GBM basin.

3.5 Comparison methods

5 Accurate and reliable observations of the lower atmosphere are crucial for numerical weather prediction (NWP) and climate change studies. RO data can be used to evaluate the accuracy of different radiosonde types (e.g., Kuo et al., 2005; Sun et al., 2010), as well as helping in identification of error sources in other satellite-based measurements (Ho et al., 2009). In this study, radiosonde observations from within the GBM
 10 basin were compared with COSMIC RO profiles for the period August 2006 to December 2013. Mean differences ($\overline{\Delta X}$) and standard deviations (SD_X) of temperature (T) and water vapour pressure (p_w) between radiosondes and RO profiles at various pressure levels were computed as (e.g., Sun et al., 2010):

$$\overline{\Delta X} = \frac{1}{n} \sum_{i=1}^n (X_i^R - X_i^C), \quad (2a)$$

$$15 \quad SD_X = \sqrt{\frac{1}{n-1} \sum_{i=1}^n \left((X_i^R - X_i^C) - \overline{\Delta X} \right)^2}, \quad (2b)$$

while for refractivity (M), the relative mean differences (Rel. $\overline{\Delta X}$) and standard deviations (Rel. SD_X) were computed as (e.g., Sun et al., 2010):

$$\overline{\Delta X} = \frac{1}{n} \sum_{i=1}^n \left(\frac{X_i^R - X_i^C}{X_i^C} \right), \quad (3a)$$

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$$SD_X = \sqrt{\frac{1}{n-1} \sum_{i=1}^n \left(\frac{(X_i^R - X_i^C)}{X_i^C} - \overline{\Delta X} \right)^2}, \quad (3b)$$

where X^R and X^C are the respective radiosonde and COSMIC temperature (T) or water vapour pressure (p_w) or refractivity (N), and $i = \{1, 2, \dots, n\}$ at each pressure level for n data points.

3.6 Interannual variability of the UTLS region

COSMIC RO mission provides an excellent opportunity to assess the interannual variability of the UTLS (subtropical) region with an average of ~ 388 RO profiles per month that are uniformly distributed across the basin since August 2006 (see, Fig. 2). The interannual variability of UTLS temperature was carried out based on the monthly COSMIC RO data and were compared with MERRA products from August 2006 to December 2013. The COSMIC RO data must be gridded for each month over the region to provide a uniform spatial pattern and to provide a fair comparison to MERRA data. The monthly gridded temperature data of COSMIC were averaged for the two regions: (a) UT (400–150 hPa or ~ 7.5 – 14.2 km) and (b) LS (70–30 hPa or ~ 18.7 – 24.0 km). The mean temperature data for the two regions were then interpolated to a spatial resolution of $0.5^\circ \times 0.5^\circ$ using the “ordinary kriging” method (details in the Appendix). The geostatistical kriging methods have been shown to be more robust and accurate than other methods such as inverse-distance-weighting, Thiessen Polygons (see, e.g., Goovaerts, 2000; Zhang and Srinivasan, 2009). The UTLS temperature for the MERRA data were obtained by averaging the layers similar to COSMIC data for the same time period. The interannual variability of UTLS temperature was assessed by removing the annual and semi-annual cycles from the time-series of COSMIC and MERRA based on the multilinear regression model. For example, if $\mathbf{X}_{n \times m}$ is a time series of COSMIC or MERRA data with $n = 89$ months and $m = 1404$ grid cells over the GBM river basin,

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the residual time series ($\hat{\mathbf{X}}_{n \times m}$) can be obtained by:

$$\hat{\mathbf{X}} = \hat{x}(l, j) = x(l, j) - \left(\hat{\beta}_2(j) \cdot \cos(2\pi t) + \hat{\beta}_3(j) \cdot \sin(2\pi t) + \hat{\beta}_4(j) \cdot \cos(4\pi t) + \hat{\beta}_5(j) \cdot \sin(4\pi t) \right), \quad (4)$$

where $x(l, j)$ represents an entry of \mathbf{X} from $l = 1, \dots, n$ and $j = 1, \dots, m$, and $\hat{\beta}_2$ to $\hat{\beta}_5$ are the regression coefficients representing the linear annual ($\hat{\beta}_2, \hat{\beta}_3$) and semi-annual ($\hat{\beta}_4, \hat{\beta}_5$) cycles derived by fitting a multilinear regression model using the least squares adjustment technique.

3.7 Tropopause temperatures and heights

The significance of tropopause as a climate indicator has been reported in a number of studies (e.g., Santer et al., 2003; Sausen and Santer, 2003; Khandu et al., 2011; Awange, 2012) and many approaches have been taken to define the tropopause layer based on various atmospheric variables (e.g., temperature, potential vorticity, etc.). This study focusses on the interannual variability of thermal tropopause or “lapse-rate-tropopause” (LRT) derived from COSMIC RO and MERRA data over the GBM river basin. According to WMO (1957), LRT is the lowest level at which lapse rate decreases to less than 2°C km^{-1} and the average lapse rate between the lowest level and anywhere above that level up to 2 km remains less than 2°C km^{-1} . The tropopause heights and temperatures from COSMIC data have already been computed by CDAAC, while MERRA provides tropopause temperature based on the WMO (1957) definition. The tropopause height, h_{LRT} (in km) for MERRA reanalysis was approximated from the tropopause pressure, p (in hPa) using the following relationship (PSAS, 2004):

$$h_{\text{LRT}} \equiv 44\,330.8 - 4946.54 \times p^{0.1902632}. \quad (5)$$

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4 Results and discussion

4.1 Comparison of radiosonde and COSMIC RO data

The quality of radiosonde types with respect to GNSS RO data has been reported in a number of studies (e.g., Rocken et al., 1997; Hajj et al., 2004; Ho et al., 2009; Sun et al., 2010; Steiner et al., 2011; Wang et al., 2013) and their overall agreement with COSMIC RO data has been found to be less than 0.15°C with a standard deviation of $1.5\text{--}2.0^{\circ}\text{C}$ (e.g., Sun et al., 2010). In this section, the quality of different radiosondes across the GBM river basin (see, Table 1) were assessed using COSMIC RO profiles between August 2006 and December 2013. Of particular interest is the performance of the recently upgraded radiosondes at three stations in India: New-Delhi, Patna, and Dirbugarh. Statistical comparisons were made with respect to temperature, water vapour pressure, and refractivity using a collocation criteria of 200 km and 2 h between COSMIC RO data and radiosonde observations. While a more closer collocation criteria was deemed necessary for comparison due to drifts in radiosondes and tangent point horizontal drifts in GNSS RO data (Sun et al., 2010; Wang et al., 2013), the more relaxed criteria was used here considering the low number of samples at some stations such as those over Bangladesh and China. The number of collocated data points varied at each pressure level for both temperature and humidity (i.e., water vapour pressure) as most radiosondes tend to malfunction or burst out before reaching the lower stratosphere as shown in Fig. 4.

4.1.1 Temperature

Figure 5 shows the mean temperature difference and their standard deviations at 12 pressure levels from 850–30 hPa (or 1.5–24 km) of various radiosonde types against COSMIC RO data over the GBM basin. Their mean difference and standard deviations in the UTLS region are presented in Table 2. The quality of radiosondes differ highly among different sonde types with the closest agreement shown by ShangE indicating

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from about 5 % (IMD) to less than 1 % (ShangE/ShangM). Thus, very large refractivity errors in IMD sondes in the upper troposphere may have resulted from anomalously warm bias (see, Fig. 5a) and possible errors in pressure measurements.

4.2 Comparison of COSMIC RO data and Numerical Weather Prediction (NWP) model outputs

To assess how radiosondes observations differ from the NWP outputs over the GBM river basin, especially in the lower GBM river basin where IMD sondes performed poorly, a brief analysis was carried out by comparing NWP outputs from ERA-Interim and GFS models, radiosonde observations, and COSMIC profiles for the year 2012.

The year 2012 was adopted arbitrarily in order to confirm the statistical results of radiosondes in the region, and thus, radiosonde observations were also included in the analysis. Mean differences and standard deviations of temperature, water vapour pressure, and refractivity were computed for the year 2012 based on the COSMIC RO data using the collocation criteria of 200 km radius and a time difference of 2 h from a radiosonde observation. Only 175 profiles were found to match this criteria, with all of them located over India and Bangladesh. First, to assess how each product differed over the individual stations, Fig. 8 plots the skew- T / \log - P over four stations from west to east of the GBM river basin: (a) Patialia (#42101), (b) Siliguri (#42397), (c) Dilbugarh (#42314), and (d) Chittagong (#41977). The skew- T / \log - P chart is a commonly used thermodynamic diagram to help forecasters determine the relative humidity, stability, vertical wind shear, severe weather, and flash flood potential of the atmosphere. In a Skew- T / \log - P chart, the y axis is the pressure (in hPa) in a log scale (black dashed lines), temperature is skewed at 45° (solid black lines), moist adiabatic (in red lines), mixing ratio (in green lines), and dry adiabatic (in blue lines).

The NWP products generally showed good agreement with the COSMIC RO data above 400 hPa (or ~ 7.32 km) but showed slightly larger difference below 400 hPa. The dry profiles of COSMIC RO data on the other hand, tend to deviate away from wet profile at around 300–400 hPa. The radiosonde observations showed very good agree-

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ment at Patialia (Fig. 8a) and Chittagong (Fig. 8d) but showed warm bias at Dilrugarh (Fig. 8c). The overall accuracy of the NWP data in the UTLS for temperature and refractivity and in the lower troposphere for water vapour pressure are provided in Table 3. The results indicate that both ERA-Interim and GFS data showed good agreement with COSMIC RO data, which far exceeded the quality of radiosonde data over the GBM river basin. For example, the mean temperature differences of ERA-Interim and GFS data were 0.15 °C with a standard deviation of around 1.0 °C while radiosonde data showed a mean difference and a standard deviation of 1.22 and 3.94 °C, respectively.

4.3 Interannual variability of the UTLS region

Figure 9 shows the seasonal variability of temperature in the UTLS region over the GBM river basin from August 2006 to December 2013. The temperature in the UT region was marked by a strong seasonal pattern indicating a difference of up to 10 °C between the lower and the upper peaks (Fig. 9a and b). The strong seasonality gradually diminishes and became nearly constant in the tropopause layer between 150 and 70 hPa (14–18.5 km) indicating a much wider tropopause belt in the region. This supports the view that tropopause is no more a thin layer (see, Fueglistaler et al., 2009). The temperature pattern in the stratosphere (above 18.5 km) was relatively stable without much seasonality but exhibited sudden fall in temperature during the early months of 2008, 2009, and 2013. These cooling events have been reported to be associated with Sudden Stratospheric Warming (SSW) events in the winter polar regions (Hansen et al., 2014). The stratosphere responds to major global climatic events (e.g., Brewer–Dobson circulation), solar radiations, and natural and man-made aerosol forcings (e.g., Fueglistaler et al., 2009), and have been found to influence global weather conditions in the lower troposphere (e.g., Foelsche et al., 2008; Seidel et al., 2011). The SSW events in the winter polar region have been found to be associated with significant cooling (of about 1 °C) of the stratosphere over low latitude and equatorial regions (Fritz and Soules, 1972).

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The MERRA data (Fig. 9b) was found to agree very well above 12.5 km indicating a difference of only $\pm 1^\circ\text{C}$ (Fig. 9c) but indicated large differences in the tropospheric region with a difference of up to $\pm 3^\circ\text{C}$. The amplitudes of COSMIC RO data were found to be lower in the lower troposphere with a slightly delayed temperature maximum in the region. The lower amplitudes in COSMIC RO (Fig. 9a) could have resulted from the low resolution background information used to retrieve RO temperature in the troposphere (Kuo et al., 2004).

To further understand the dynamics of the UTLS region over the GBM basin, the interannual variability of temperature of the UTLS region was assessed. The temporal temperature patterns were derived by removing the annual and semi-annual cycles from the basin-averaged time series of COSMIC RO and MERRA data (see, Eq. 4). The de-seasonalised time series were compared with large-scale ocean-atmospheric phenomena such as ENSO and IOD, which were found to significantly influence the seasonal climate fluctuations in the GBM river basin (see, e.g., Chowdhury and Ward, 2004; Ashok and Saji, 2007). ENSO is commonly represented by sea surface temperature (SST) anomalies such as those within the Niño3.4 region (5°N – 5°S , 120 – 170°W) (Trenberth, 1990), while IOD is commonly measured by the difference between SST anomalies in the western (50 – 70°E and 10°S – 10°N) and eastern (90 – 110°E and 10 – 0°S) equatorial Indian Ocean, which is referred to as Dipole Mode Index (Saji et al., 1999). Monthly ENSO index (Niño3.4) from the Climate Prediction Center (CPC, http://www.cpc.ncep.noaa.gov/products/analysis_monitoring/ensostuff/) and DMI (as IOD index) from the Low-latitude Climate Prediction Research (http://www.jamstec.go.jp/frcgc/research/d1/iod/iod/dipole_mode_index.html) were used.

In addition, comparisons were also made with Quasi-Biennial Oscillation (QBO) index at 30 hPa, which were extracted from the NOAA website (<http://www.esrl.noaa.gov/psd/data/climateindices/list/>). QBO or formally known as Singapore Winds is a quasiperiodic oscillation of the equatorial zonal wind between the easterlies and westerlies in the tropical stratosphere where the downward propagating winds are finally dissipated at the tropical tropopause region, with a mean period of about 28 months (Baldwin

et al., 2001). The variability of QBO in the lower stratospheric temperature can induce interannual variability in the tropopause temperature (Baldwin et al., 2001). QBO is measured by an index derived from the zonally averaged easterlies at 50–30 hPa above the equatorial region.

Figure 10a and b shows the interannual variability of temperature in the UT and LS, respectively, together with the corresponding temporal patterns of QBO, ENSO (Niño3.4), and IOD (DMI) over the period from August 2006 to December 2013. The temporal patterns of temperature in UT and LS region were found to exhibit the patterns of all three climate indices with varying degree of magnitudes. The ENSO activity heightened during the past 5 years with two major warm phases (i.e., El Niño or +ive Niño3.4) in 2006/07 and 2009/10 and four major cold phases (i.e., La Niña or -ive Niño3.4) during the periods 2007/08, 2008/09, 2010/11, and 2011/12 (see, Fig. 10). These cold (warm) ENSO phases are associated with extreme drought conditions (flooding) in the GBM river basin with further contribution from IOD activities, which may also act independently during the neutral ENSO periods (e.g., Chowdhury, 2003; Ashok and Saji, 2007).

The stratospheric temperature showed a maximum amplitude of -3°C during the winter of 2008/09 (Fig. 10a) corresponding to the major SSW event (e.g., Resmi et al., 2013; Hansen et al., 2014) whereas the maximum amplitude (3°C) of tropospheric temperature was found to occur in the winter of 2011/12 corresponding to a moderate La Niña period (Fig. 10b). Similarly, the stratospheric cooling in the winter of 2012/13 were associated with a major SSW event (see, <http://gmao.gsfc.nasa.gov/researchhighlights/SSW/>) but has negligible impact on the tropospheric variability. The stratospheric temperature increases (decreases) during the positive El Niño (La Niña) periods (Fig. 10a) while the tropospheric temperature showed the opposite, indicating an increase in temperature during the warm ENSO phase (Fig. 10b). Thus, both stratospheric and tropospheric temperature shows very good correlation with ENSO pattern with a magnitude of 0.44 (for stratospheric temperatures) and -0.62 (for tropospheric temperatures) based on the COSMIC RO data even for a very short period of time

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(see, Table 4). The MERRA data showed a slightly better correlation with ENSO but indicated a lag of approximately one month with respect to the tropospheric temperature. This lag can be found especially during the periods of 2008/09, 2010/11, and 2013 but generally agrees very well with COSMIC RO data.

The relationship between IOD and UTLS was found to follow the ENSO pattern indicating positive correlation during the warm ENSO phases and negative correlation during the cold phases of ENSO (see, Fig. 10). The correlation between IOD and tropospheric temperature based on the 89 months of COSMIC RO data was -0.59 with a lag of 7 months (see, Table 4). Similar results were shown by MERRA data for the same lag period. Since QBO is a stratospheric phenomenon, it has greater influence on the stratospheric temperature than on the tropospheric temperature. As a result, QBO has been found to be highly correlated with the stratospheric temperature (0.5 to -0.5).

4.4 Tropopause heights and temperatures over GBM basin

The seasonal and interannual variations of tropopause temperatures and heights over the GBM river basin were investigated. Figure 11 shows the frequency distribution of tropopause temperatures and heights based on the COSMIC RO data only. The tropopause exhibited a clear seasonal variation over the GBM basin with lower and warmer tropopause during winter (DJF, i.e., Fig. 11a and e), and higher and colder tropopause during the summer (JJA, i.e., Fig. 11c) and autumn (SON, i.e., Fig. 11d). The tropopause height varied between 10 km and a little over 18.5 km corresponding to a temperature of -40°C and around -85°C . Large variations were seen especially during the spring (MAM) with a standard deviation of 9.6°C and 2.2 km, respectively (Fig. 11b and f). This strong variation of tropopause in spring could be due to high diurnal (day and night) variations, which tend to be in the order of ~ 1 km and $\sim 3^{\circ}\text{C}$ over the region (Mehta et al., 2010).

To further illustrate the seasonal variations of the tropopause, Figs. 12 and 13 shows the spatial distribution of annual and semi-annual amplitudes of tropopause temper-

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atures and heights (i.e., mean annual variation) derived from MERRA and COSMIC RO data. These annual and semi-annual amplitudes were derived from the multilinear regression model described in Eq. (4). The highest amplitudes were found over the subtropics (north of 26°) where the tropopause temperatures were found to vary by almost 8°C (Fig. 12a and c) and tropopause heights varied by up to 2 km in the same region (Fig. 13a and c). The semi-annual variations of tropopause temperatures (Fig. 12b and d) and heights (Fig. 13b and d) were relatively small (with a maximum of 0.4°C and 0.15 km, respectively). It is interesting to note that the magnitude of variations in tropopause temperatures increased sharply along 26°N (Fig. 12a and c) while the tropopause heights' variations were found to be more steeper along northern boundary of the GBM river basin. The spatial patterns of amplitudes derived from MERRA data as shown in Figs. 12c and 13c were found to be consistent with COSMIC RO data, but MERRA indicated larger temperature variations while its height variations were less than 1.5 km.

To be more precise on the seasonal variation of tropopause over the region, the mean annual cycle of tropopause temperatures and heights with their standard deviations are plotted in Fig. 14. The seasonal variation of tropopause indicated a strong annual cycle following the Indian monsoon over the GBM river basin. The minimum tropopause temperature of COSMIC RO data was found to occur during the peak monsoon period (July) while the maximum temperature was found to occur in January (Fig. 14a). On the other hand, the tropopause height peaks in June and reaches its minimum in January (Fig. 14b). The tropopause variations (represented by error bars) were found to be very small during the monsoon period (June–September) and substantially high from November to May. Such variations could be related to synoptic scale (day-to-day) variability of the regional atmospheric convection over the region (Mehta et al., 2010). The day-to-day variations can be particularly large when the small-scale convective activities dominate the regional atmosphere between November and May. Both the dataset indicated strong annual cycle but tropopause was found to be warmer (by $2\text{--}5^\circ\text{C}$) and lower (by 1 km) over the GBM basin. Earlier studies have

also reported warm biases in reanalysis products compared to GNSS RO data and radiosonde observations (e.g., Randel et al., 2000).

While it is impossible to derive long-term trends and variability of the tropopause using COSMIC RO data, a short-term (89 months) interannual variability of the tropopause temperatures and heights has been carried out here. Since tropopause is a transitional layer between the troposphere and the stratosphere, both tropospheric activities such as ENSO and IOD and stratospheric activities such as QBO and SSW events can have significant influence on the temporal variations of tropopause heights and temperatures (Randel et al., 2000). The de-seasonalised temporal anomalies (using Eq. 4) (i.e., deviation from mean) of tropopause temperatures and heights of COSMIC RO data have been compared with those from MERRA data to assess the temporal variability of tropopause over the region.

Figure 15 shows the interannual variability of tropopause temperatures and heights derived from COSMIC RO and MERRA data for the period between August 2006 and December 2013, as well as the indices of QBO, ENSO (Niño3.4), and IOD (DMI). The tropopause temperature of COSMIC RO data exhibited the patterns of ENSO (Niño3.4) and IOD (-DMI) very well (Fig. 15a) with a correlation of 0.65 and -0.38 , respectively, which were equally represented by the MERRA data (Table 5). While ENSO lags tropopause temperature by one month, IOD was found to appear earlier by 6 months as indicated by both the dataset. For instance, the tropopause warmed by about 2°C during the major El Niño event of 2009/10 and a cooling of similar magnitude has occurred during the major La Niña period of 2010/11 (see, Fig. 15a).

While the tropopause cools as a result of major SSW events in the northern polar region, even though with a lesser magnitude in this study, the major SSW event of 2009 was found to have a large impact on the tropopause heights (Fig. 15b). The tropopause height increased by more than 2 km over the region as indicated by the MERRA data. However, COSMIC RO data showed an increase of only about 1 km. ENSO and IOD events were still found to dominate the temporal variation of tropopause heights in the region, which showed a correlation of -0.52 and 0.33 (based on COSMIC RO data)

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and lag (lead) of one (six) months, respectively. MERRA data indicated a slightly higher correlation with ENSO and IOD. The increase in tropopause temperature during the El Niño conditions were associated with an equivalent decrease in tropopause heights (e.g., 2009/10). Furthermore, the tropopause heights showed a similar decrease during the early months of 2013 after a moderate El Niño event (with contributions from IOD).

The QBO is a well-known stratospheric phenomenon, which has been reported to have greater influence on the tropical tropopause characteristics (e.g., Reid and Gage, 1985; Randel et al., 2000). The relationship between tropopause and QBO signal over the GBM river basin (in the sub-tropical region) was found to be quite complex mainly because of the ~ 28 month cycle of the QBO signal (Baldwin et al., 2001), which cannot be properly captured in the temporal period defined in this study. The peak-to-peak time differences of maximum correlation (0.4 to -0.5) between tropopause (temperatures and heights) and QBO signals was around 30 months, consistent with the cycle of QBO signal (figures not shown).

5 Conclusions

Several GNSS RO missions were launched in the past decade to provide high-quality, global observations of the Earth's atmosphere (ionosphere, stratosphere, and troposphere) for various weather and climate applications. This study examined the interannual variability of temperature in the upper troposphere/lower stratosphere (UTLS) region as well as the tropopause over the GBM river basin in South Asia using 89 months (August 2006 to December 2013) of COSMIC RO data. The GBM basin received a total of 35 776 COSMIC RO profiles over the study period, with an average of ~ 388 well-distributed profiles/month from August 2006–December 2013. More than 56 % of the profiles reached at least 1.5 km above the mean sea level height. Radiosonde observations from 24 stations in and around the GBM river basin were compared with COSMIC RO data to assess the quality of on-going monitoring work in the region. ShangE/China radiosondes showed the best agreement with a mean temperature dif-

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ference of -0.06°C and standard deviation of 1.44°C followed by ShangM/China and those over Bangladesh. The IMD radiosondes still exhibited large warm biases and anomalously high variations in the UTLS region even after upgrading their radiosondes at three stations. However, from the data period used in this study, the mean temperature differences were observed to have reduced by more than half (2.72 to 1.20°C) after upgrading.

The temporal anomalies of UTLS (400–30 hPa; 7.5–24.0 km) temperatures exhibited clear temporal patterns of ENOS/IOD modes while also adequately capturing the major sudden stratospheric warming (SSW) events of 2009 and 2013. Overall, the tropospheric temperature anomalies indicated very good correspondence with ENSO (Niño3.4 index) and IOD (DMI index) with negative correlations of 0.62 and 0.59, respectively. The MERRA data performed equally well with even higher correlation in this study. Based on COSMIC RO data, the mean tropopause temperature was found to vary between -70 to -80°C with an average height of about 15.5 to 16.3 km from winter to summer over the GBM river basin, thus, exhibiting a strong annual cycle in the region. The annual amplitudes of tropopause temperatures varied considerably from the tropics to the sub-tropics (0 to 5°C) while the corresponding tropopause heights varied from 0 to 1.5 km.

Both tropopause temperature and height anomalies over the GBM river basin exhibited strong temporal patterns of ENSO indicating a correlation of 0.65 and -0.52 , respectively, with a lag of one month, and were equally represented by the MERRA data. The IOD also indicated reasonably good correspondence with tropopause temperatures and heights with a correlation of -0.38 and 0.33, and with a lag of six and five months, respectively. While QBO has a well-known interannual signal, which significantly impacts the variability of tropopause in the tropics, its correlation with tropopause over the GBM river basin varied between 0.4 to -0.5 with a peak-to-peak time difference of ~ 30 months. The tropopause heights were found to be sensitive to the recent SSW events of 2008/09, but were not properly captured during the last two SSW events

opposed to a complex terrain and its associate weather in the region. Based on our experiments we have adopted the spherical model for all the months.

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Table 4. Correlation coefficients between UTLS temperature and climate indices between August 2006 to December 2013.

Climate Index	Stratospheric Temperature		Tropospheric Temperature	
	COSMIC	MERRA	COSMIC	MERRA
Nino3.4	0.44 (0)	0.54 (1)	−0.62 (1)	−0.67 (1)
DMI	0.26 (0)	0.39 (1)	−0.59 (−7)	−0.56 (−7)

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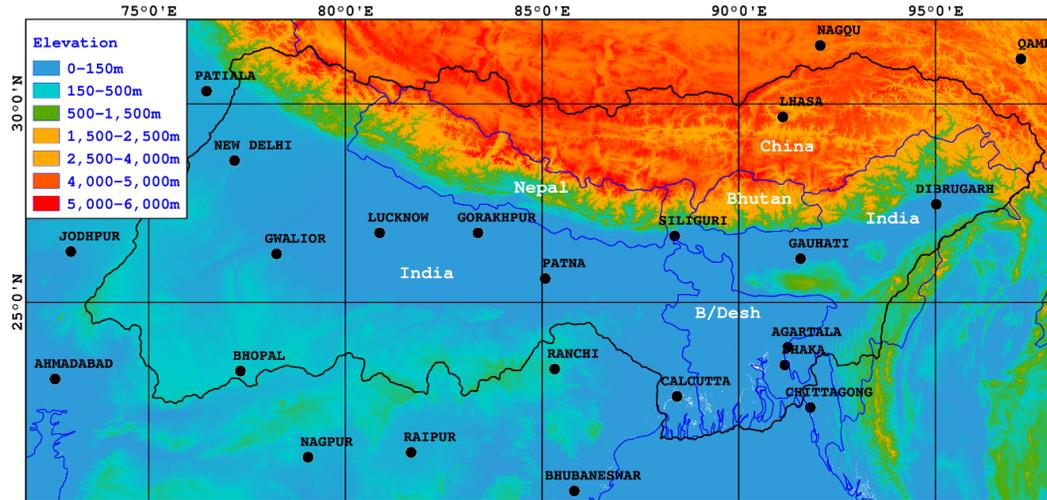


Figure 1. Elevation of the Ganges–Brahmaputra–Meghna basin in South Asia. The digital elevation model is derived from the Shuttle Radar Topography mission (SRTM, <http://srtm.csi.cgiar.org>). The locations of the existing radiosonde stations are shown in circles (black).

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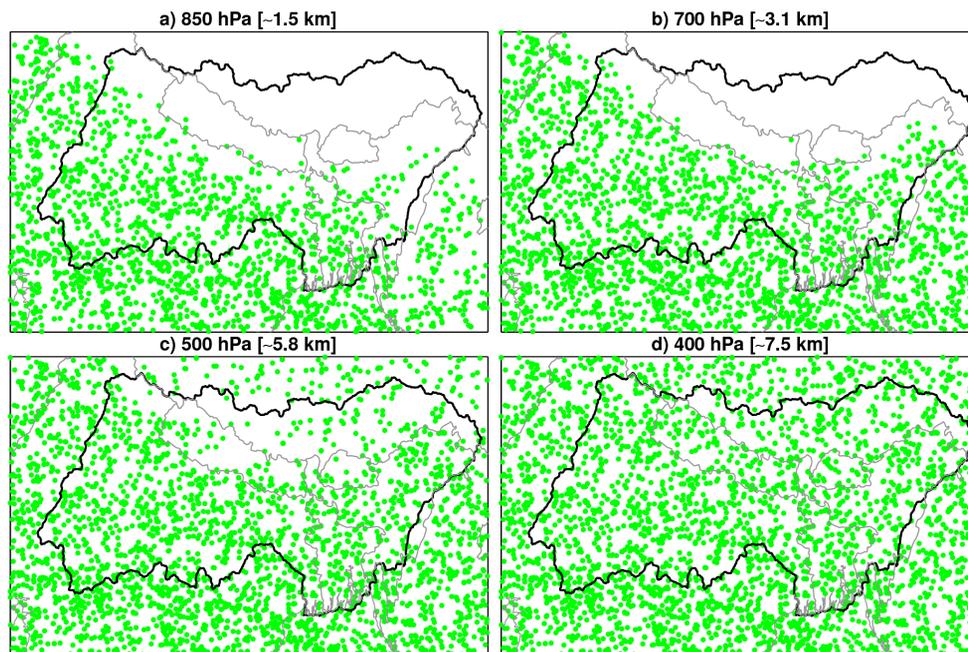


Figure 3. (a) Spatial distribution of COSMIC data points in the lower troposphere for the year 2012: (a) 850 hPa (~ 1.5 km), (b) 700 hPa (~ 3.1 km), (c) 500 hPa (~ 5.8 km), and (d) 400 hPa (~ 7.5 km).

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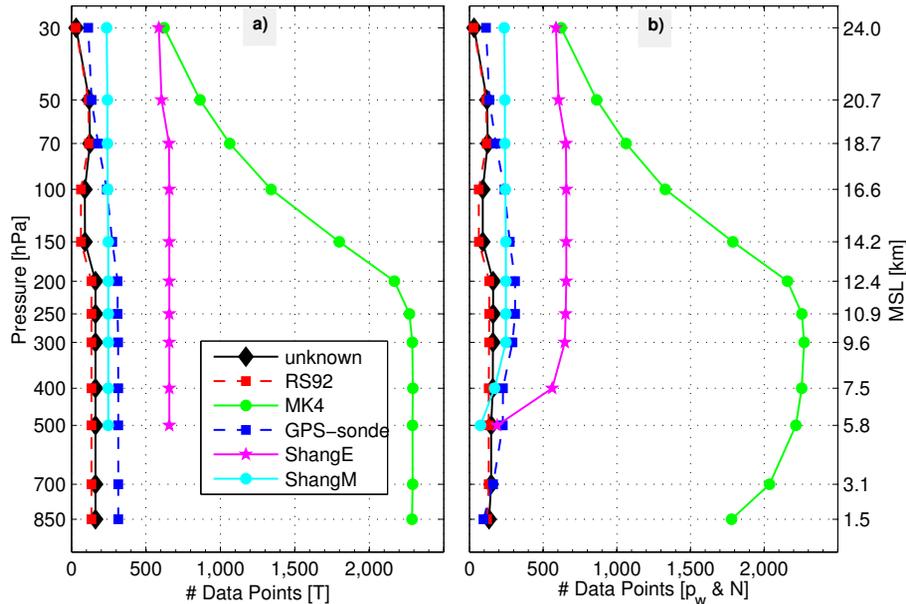


Figure 4. Number of data points at each pressure levels for temperature and humidity (water vapour pressure) of all the radiosonde types (see, Table 1) from August 2006 to December 2013.

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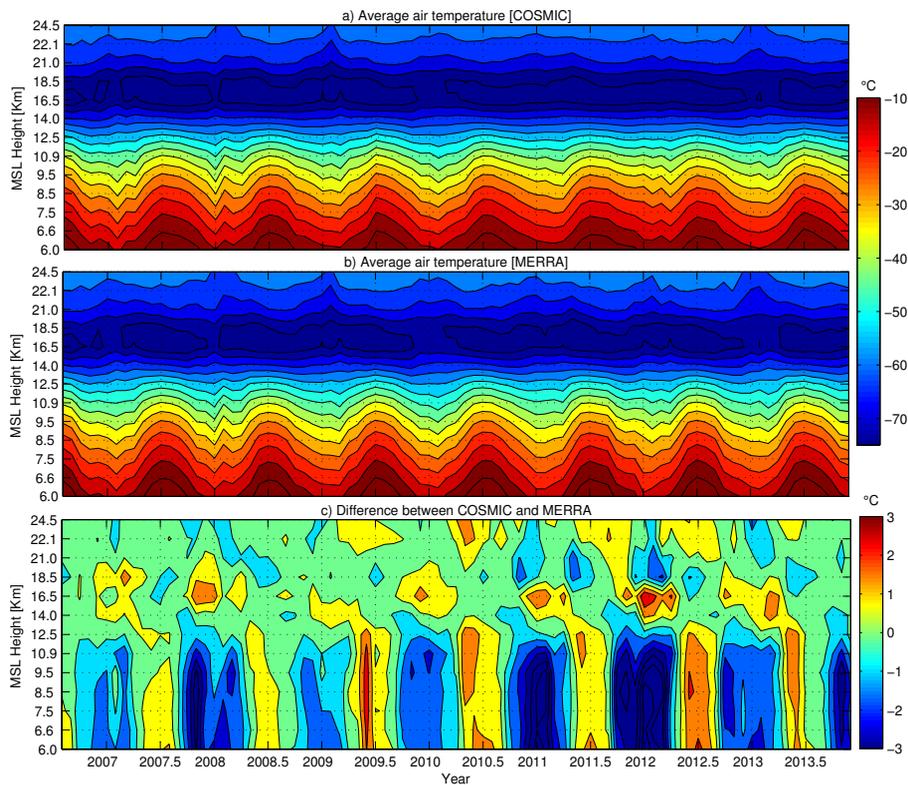


Figure 9. Mean air temperature between 500–30 hPa (or 6.0–24.5 km) for the period August 2006– December 2013. **(a)** COSMIC wet profiles, **(b)** MERRA re-analysis product, and **(c)** mean temperature difference between MERRA and COSMIC data.

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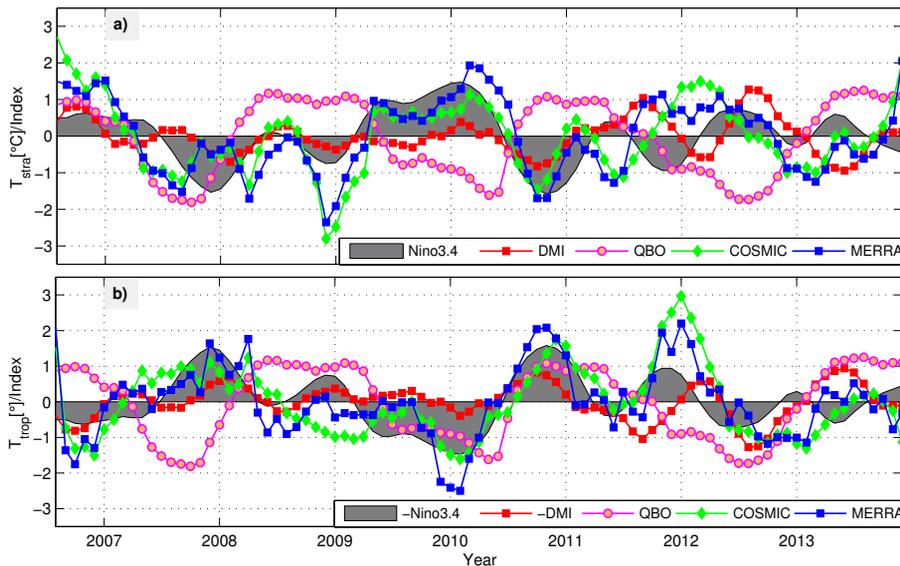


Figure 10. Interannual temperature variability of **(a)** LS (70–30 hPa; 18.5–24.5 km) and **(b)** UT (400–150 hPa; 7.5–14.0 km) based on MERRA data and COSMIC RO data from August 2006 to December 2013. QBO, ENSO and IOD indices are also shown. Note that **(b)** plots La Niña hence the opposite **(a)**.

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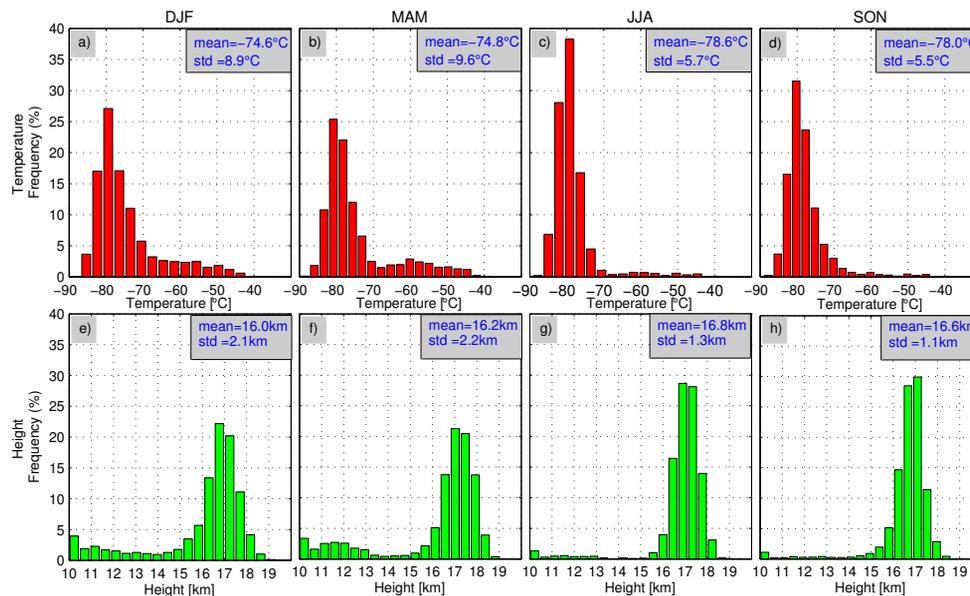


Figure 11. Frequency distribution of tropopause temperatures (a–d) and heights (e–h) over the GBM basin for the period between August 2006 and December 2013 based on COSMIC RO data.

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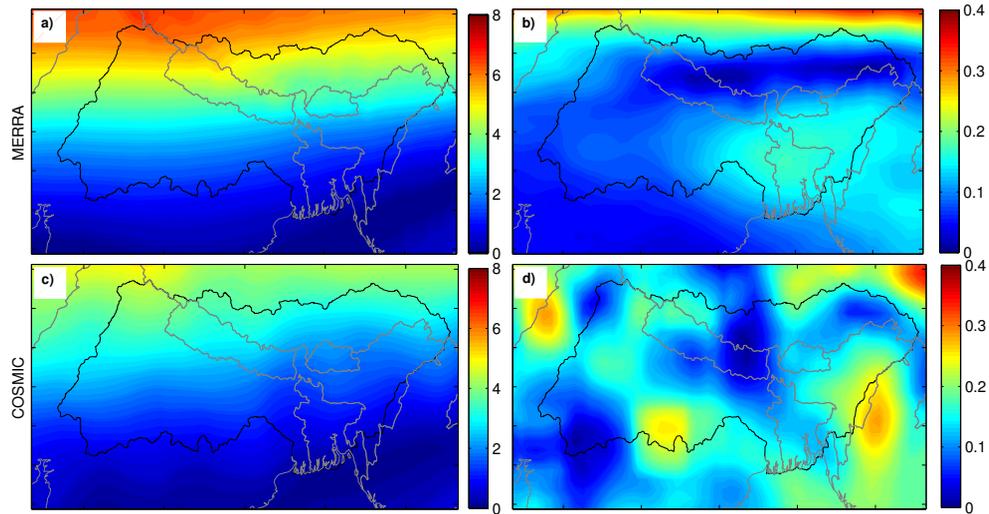


Figure 12. Annual and semi-annual amplitudes of tropopause temperatures (in °C) derived from MERRA (a and b) and COSMIC RO (c and d) data from August 2006 to December 2013. The annual amplitudes are scaled between 0–8 °C (a and c) and semi-annual amplitudes are scaled between 0–0.4 °C (b and d).

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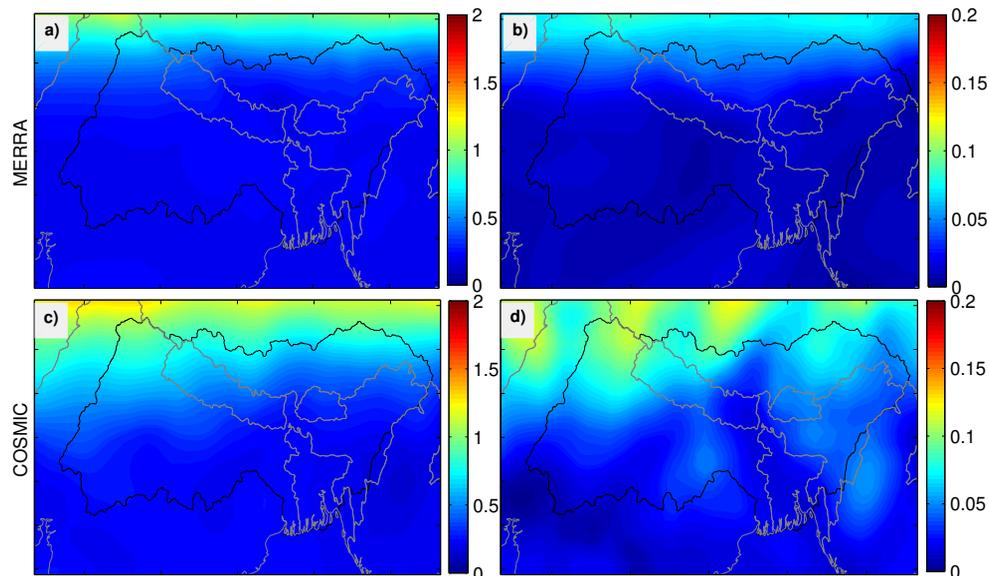


Figure 13. Annual and semi-annual amplitudes of tropopause heights (in km) derived from MERRA (a and b) and COSMIC RO (c and d) data from August 2006 to December 2013. The annual amplitudes are scaled between 0–2 km (a and c) and semi-annual amplitudes are scaled between 0–0.2 km (b and d).

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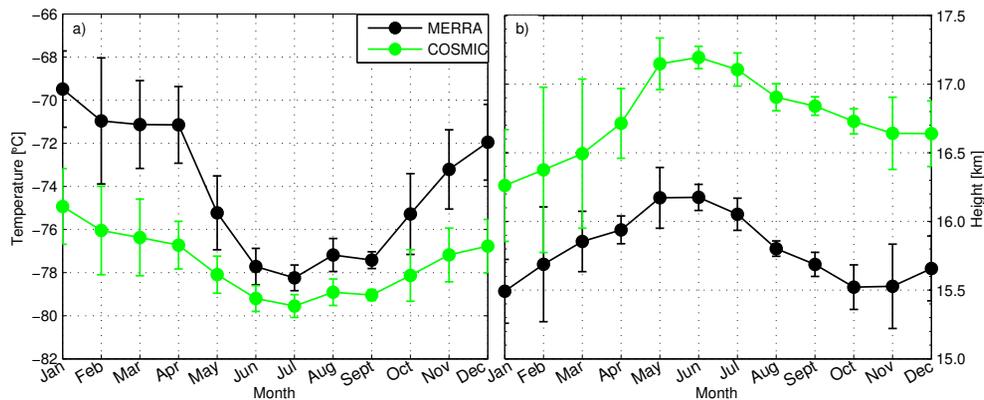


Figure 14. Annual cycle of tropopause over the GBM river basin computed from MERRA and COSMIC RO data for the period between August 2006 and December 2013: **(a)** tropopause temperatures, and **(b)** tropopause heights. The seasonal variations correspond reasonably well with Figs. 12 and 13 considering that factors impacting the interannual variabilities were effectively removed in Figs. 12 and 13 based on the multi-linear regression approach in Eq. (4).

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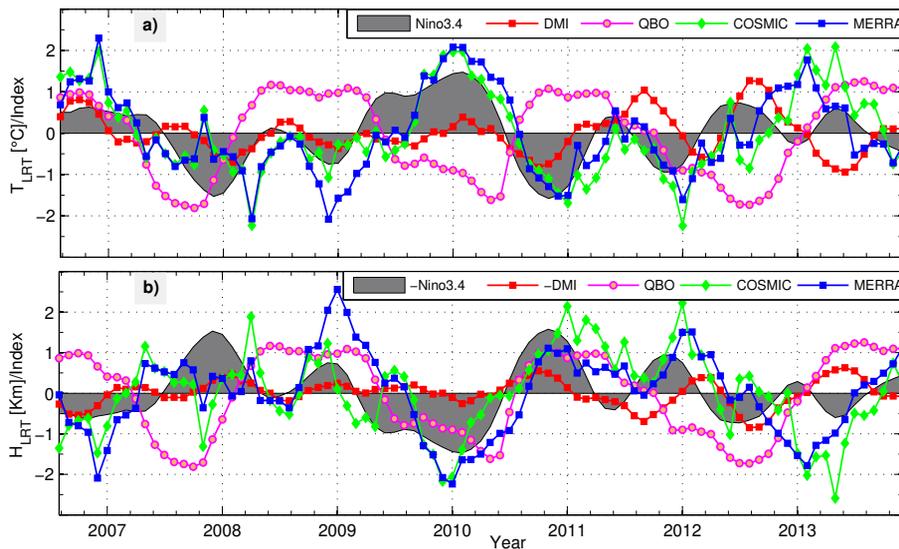


Figure 15. Interannual variability of tropopause temperatures and heights over GBM basin based on MERRA data and COSMIC RO data for the period August 2006 to December 2013, and their relations to major climatic indices in the region.

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